

時系列解析（3）

配布資料（2018/4/25）

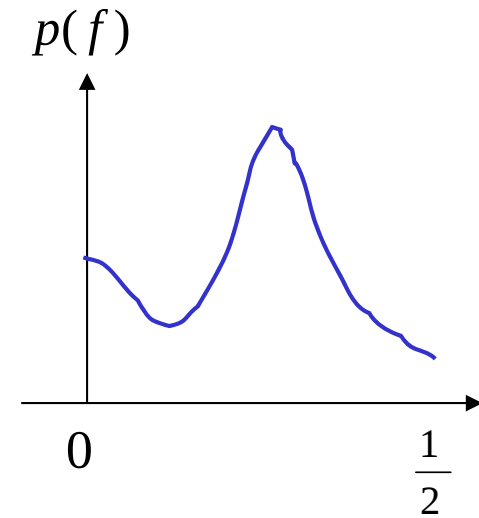
東京大学 数理・情報教育研究センター
北川 源四郎

(パワー) スペクトル

$$C_k \quad \text{自己共分散関数} \quad \sum_{k=-\infty}^{\infty} |C_k| < \infty$$

$$p(f) = \sum_{k=-\infty}^{\infty} C_k e^{-2\pi i k f} \quad -\frac{1}{2} \leq f \leq \frac{1}{2}$$
$$= C_0 + 2 \sum_{k=1}^{\infty} C_k \cos 2\pi k f$$

$$p(-f) = p(f)$$

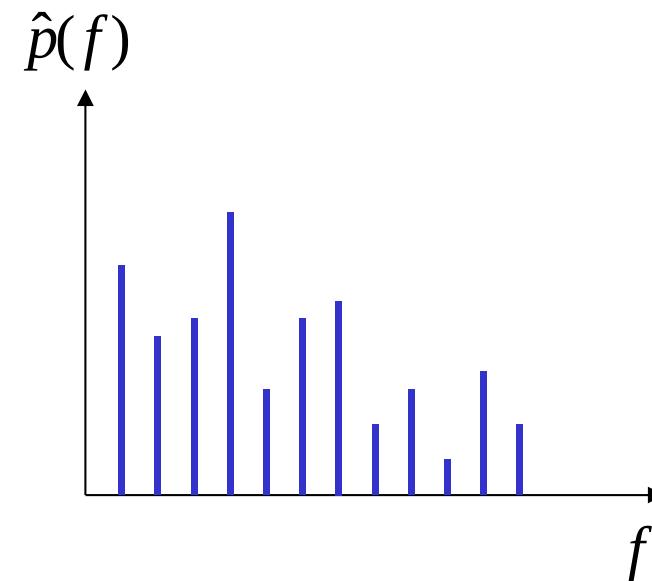


$$C_k = \int_{-\frac{1}{2}}^{\frac{1}{2}} p(f) e^{2\pi i k f} df = \int_{-\frac{1}{2}}^{\frac{1}{2}} p(f) \cos(2\pi k f) df$$

ピリオドグラム

$$y_1, \dots, y_N \quad \rightarrow \quad \hat{C}_0, \dots, \hat{C}_{N-1}$$

$$\begin{aligned} \hat{p}(f) &= \sum_{k=-N+1}^{N-1} \hat{C}_k e^{-2\pi i k f} \\ &= \hat{C}_0 + 2 \sum_{k=1}^{N-1} \hat{C}_k \cos(2\pi k f) \\ f &= 0, \frac{1}{N}, \frac{2}{N}, \dots, \frac{1}{2} \end{aligned}$$



$$\hat{C}_k = \int_{-\frac{1}{2}}^{\frac{1}{2}} \hat{p}(f) e^{2\pi i k f} df \quad k = 0, 1, \dots, N-1$$

$$\hat{p}(f) = \sum_{k=-N+1}^{N-1} \hat{C}_k e^{-2\pi i k f} \quad -\frac{1}{2} \leq f \leq \frac{1}{2}$$

ピリオドグラムの性質

(1) 漸近的に不偏

$$\lim_{n \rightarrow \infty} E[\hat{p}(f)] = p(f)$$

(2) 一致推定量でない

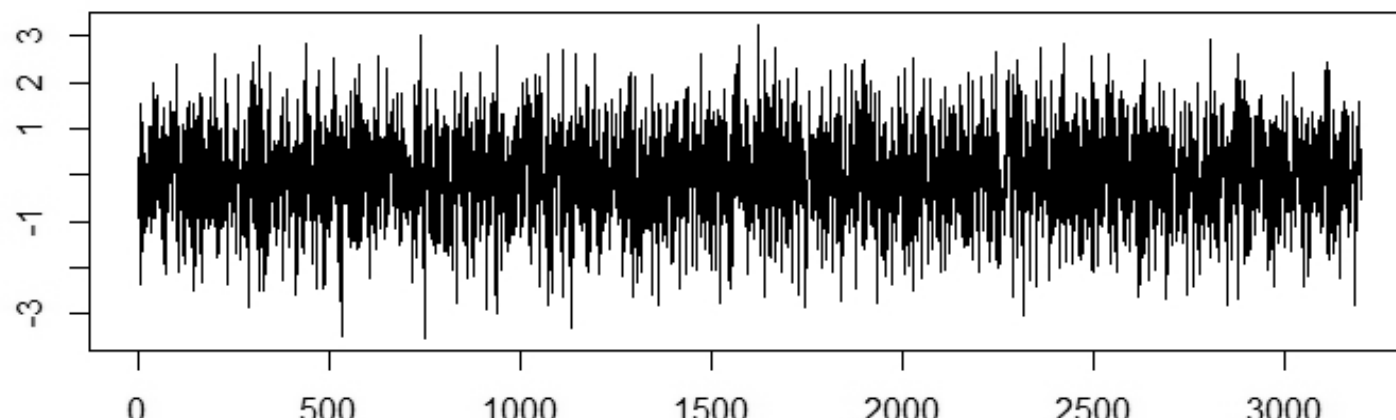
$$\lim_{n \rightarrow \infty} \hat{p}(f) \neq p(f)$$

$$\frac{2\hat{p}(f_1)}{p(f_1)}, \dots, \frac{2\hat{p}(f_m)}{p(f_m)} \sim \chi^2_2 \quad m = \left[\frac{N}{2} \right] - 1$$

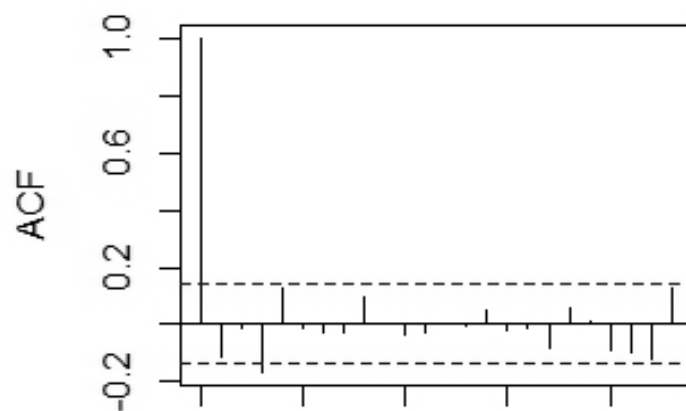
$$\frac{\hat{p}(0)}{p(0)}, \frac{\hat{p}(\frac{1}{2})}{p(\frac{1}{2})} \sim \chi^2_1$$

```
r <- rnorm(3200)
```

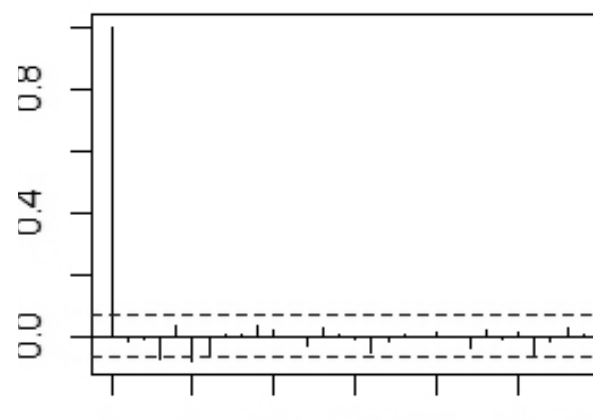
r



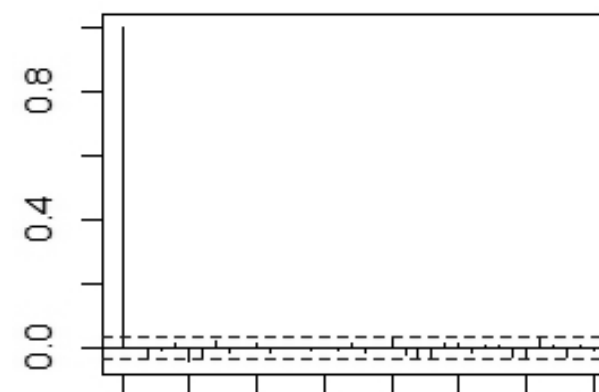
Series r[1:200]



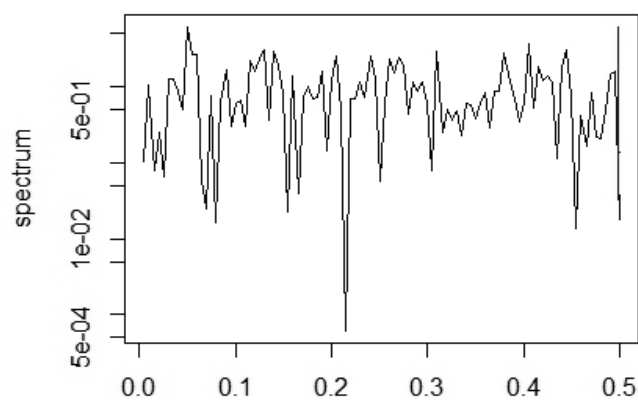
Series r[1:800]



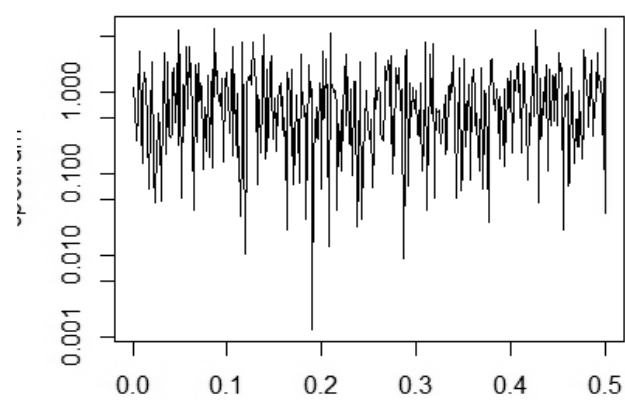
Series r



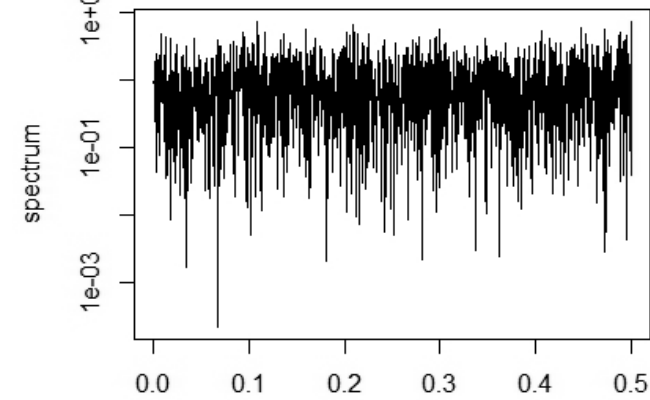
Series: r[1:200]
Raw Periodogram



Series: r[1:800]
Raw Periodogram



Series: r
Raw Periodogram



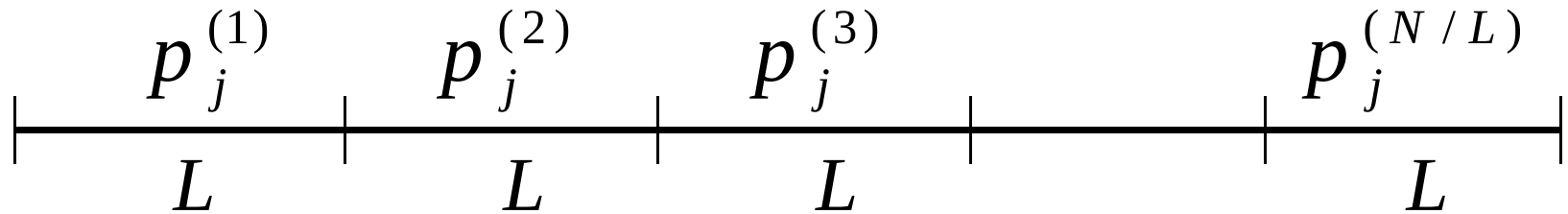
frequency
bandwidth = 0.00144

frequency
bandwidth = 0.000361

frequency
bandwidth = 9.02e-05

ピリオドogramの平均

$$\ell = \frac{N}{L}$$



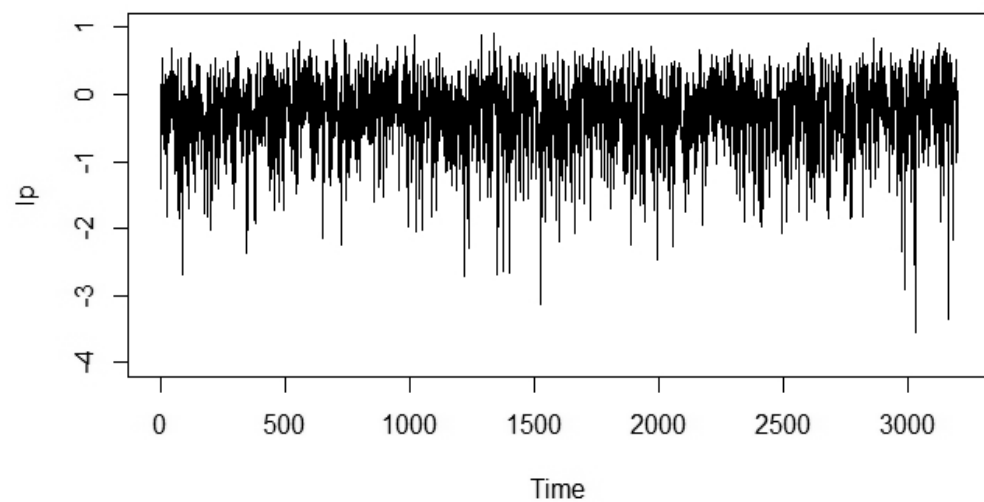
$$\hat{p}_j = \sum_{k=1}^{N/L} p_j^{(k)}$$

分散は $1/\ell$ に減少

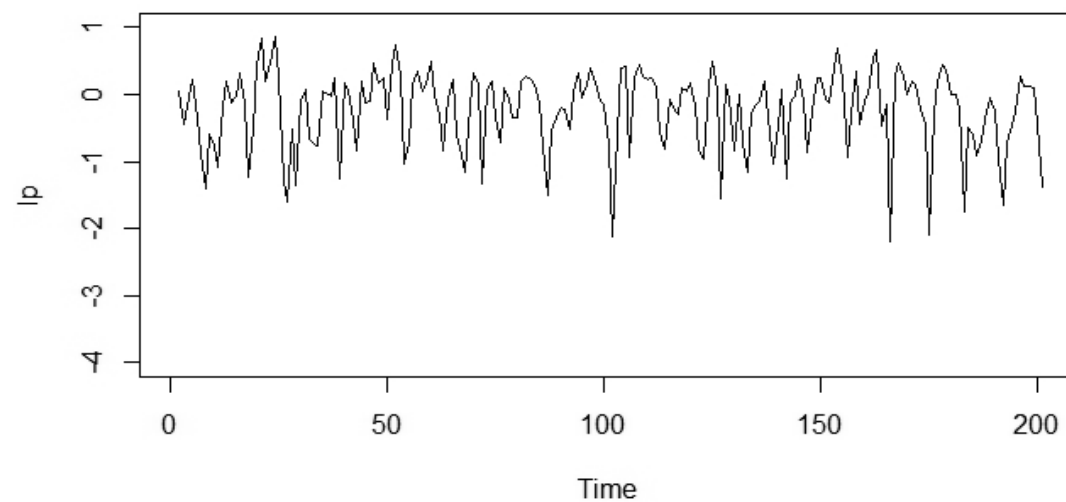
実はラグ $L-1$ までのフーリエ変換でよい

$$p_j = \hat{C}_0 + 2 \sum_{k=1}^{L-1} \hat{C}_k \cos(2\pi k f_j)$$

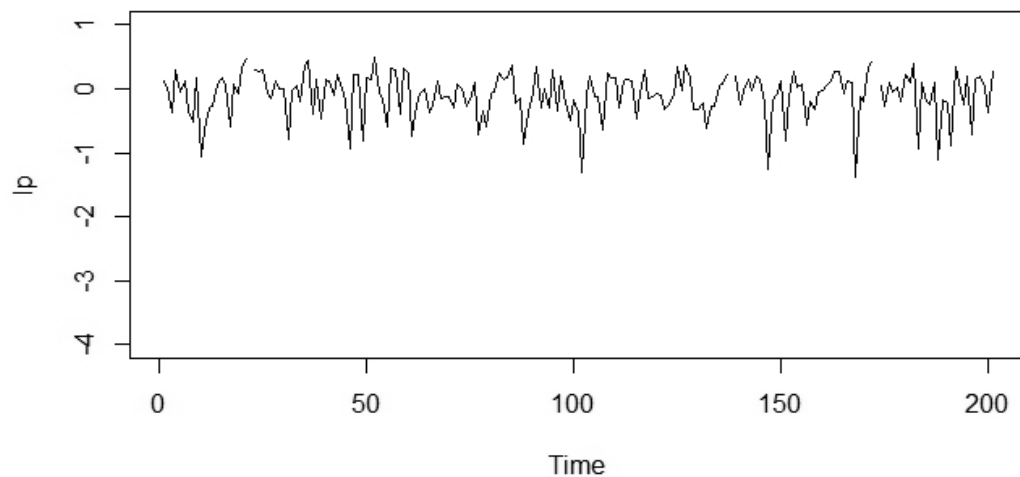
$N=3200, \text{lag}=3200$



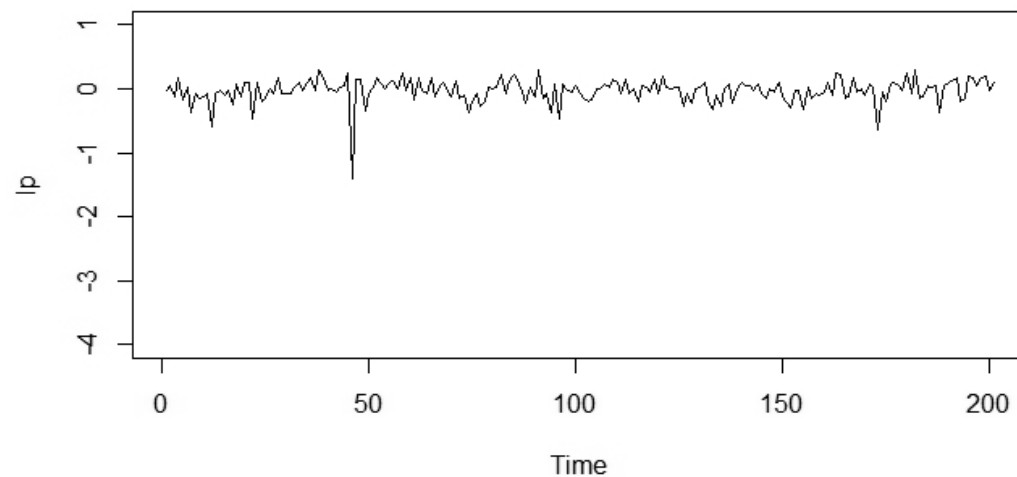
$N=200, \text{lag}=200$



$N=800, \text{lag}=200$



$N=3200, \text{lag}=200$



ピリオドグラムの平滑化

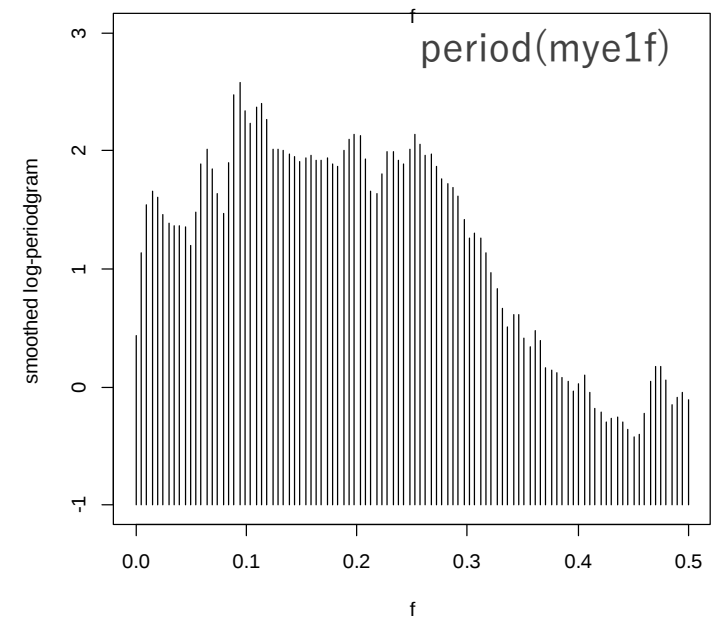
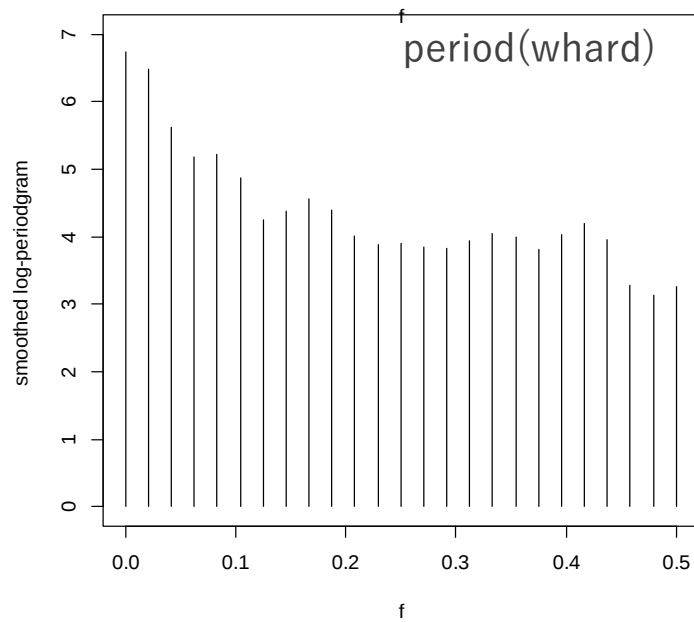
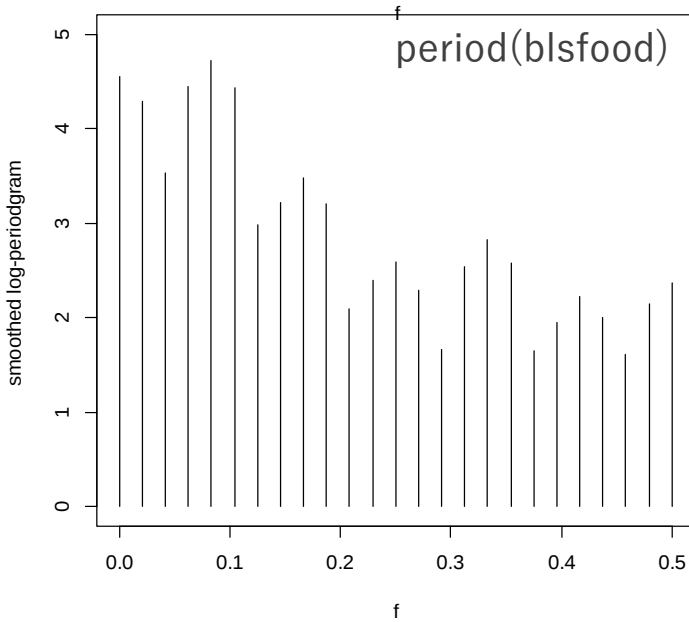
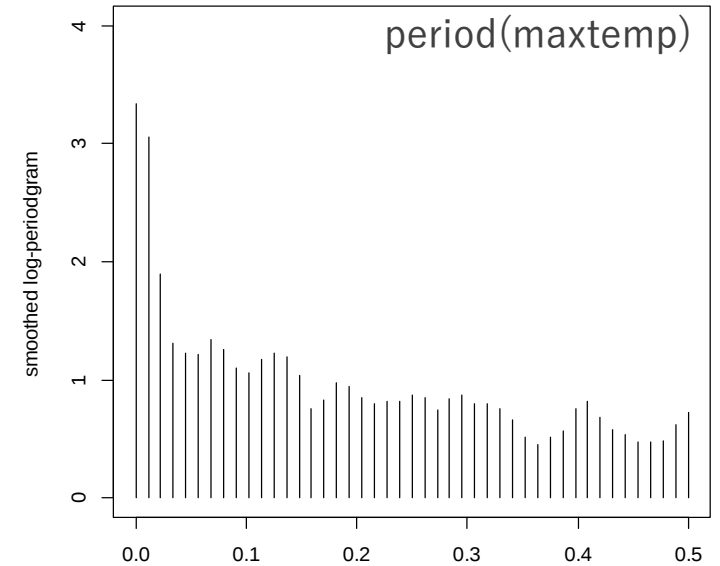
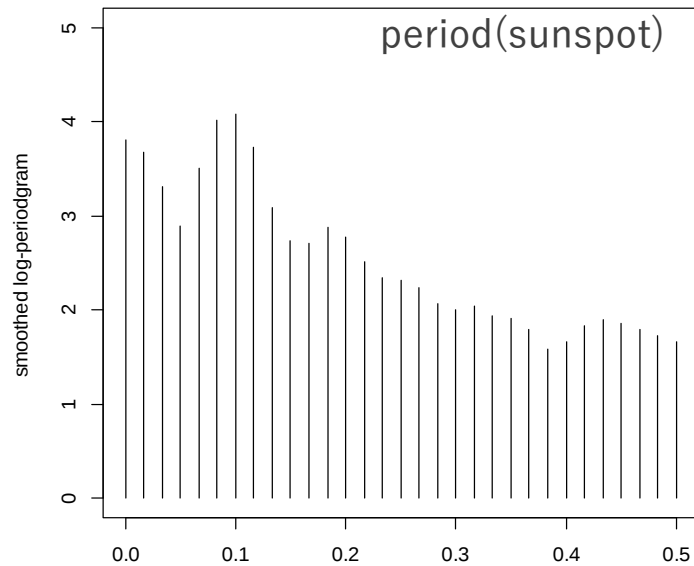
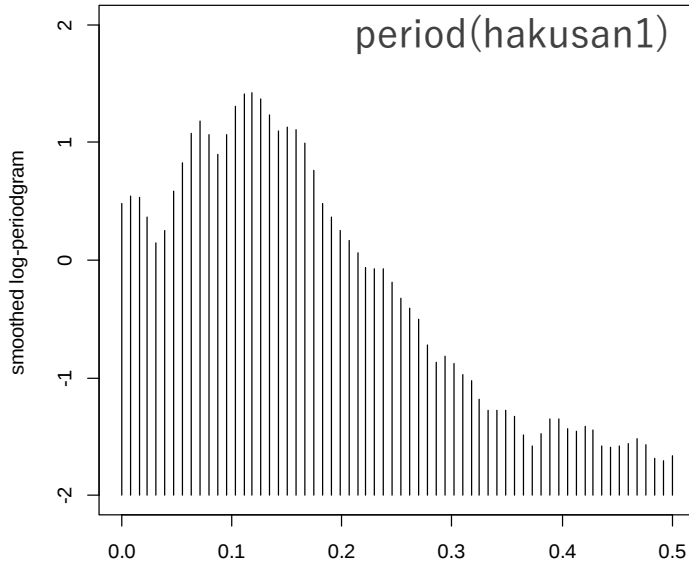
$$\tilde{p}_j = \sum_{i=-m}^m W_i \hat{p}_{j-i}$$

W_i : スペクトル ウィンドウ

Window	m	W_0	W_1
Hanning	1	0.50	0.25
Hamming	1	0.54	-.23

Blackman-Tukey 法

Smoothed Periodogram



FFT (Fast Fourier Transform)

$N = p^\ell$ のとき, 計算量 $N^2 \rightarrow Np^\ell$

$$\frac{Np^\ell}{N^2} = \frac{p^\ell}{N}$$

$$N = 1024 = 2^{10} \quad \frac{2 \times 10}{1024} \sim \frac{1}{50}$$

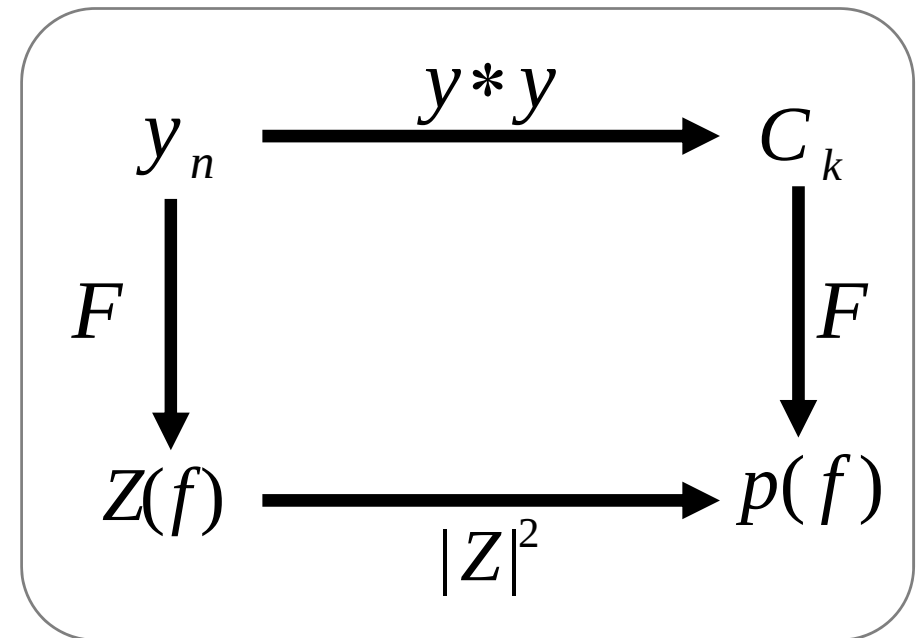
$$N = 4096 = 2^{12} \quad \frac{2 \times 12}{4096} \sim \frac{1}{170}$$

ただし $\hat{C}_k$ の計算も $O(N^2)$

FFTによるピリオドグラムの計算

$$\begin{aligned} Z_j &= \sum_{n=1}^N y_n e^{-\frac{2\pi i(n-1)j}{N}} \\ &= \sum_{n=1}^N y_n \cos\left(\frac{2\pi(n-1)j}{N}\right) - i \sum_{n=1}^N y_n \sin\left(\frac{2\pi(n-1)j}{N}\right) \\ &= FC_j - iFS_j \end{aligned}$$

$$\begin{aligned} p_j &= \frac{|Z_j|^2}{N} = \frac{1}{N} \left| \sum_{n=1}^N y_n e^{-2\pi i(n-1)j/N} \right|^2 \\ &= \frac{1}{N} (FC_j^2 + FS_j^2) \end{aligned}$$



```
# simulation by 2 cosine function
```

```
t <- 1:400
```

```
r <- rnorm(400)
```

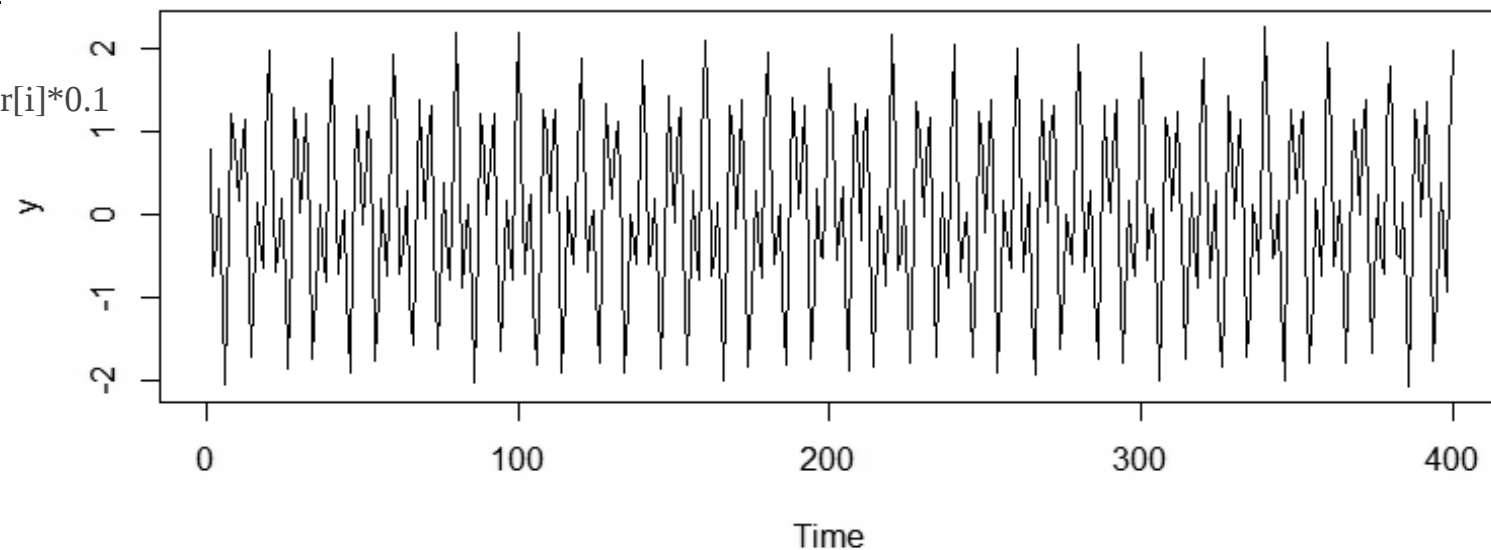
```
for (i in t) {
```

```
y[i] <- cos(2*pi*i/10) + cos(2*pi*i/4) + r[i]*0.1
```

```
}
```

```
y <- as.ts(y)
```

```
plot(y)
```



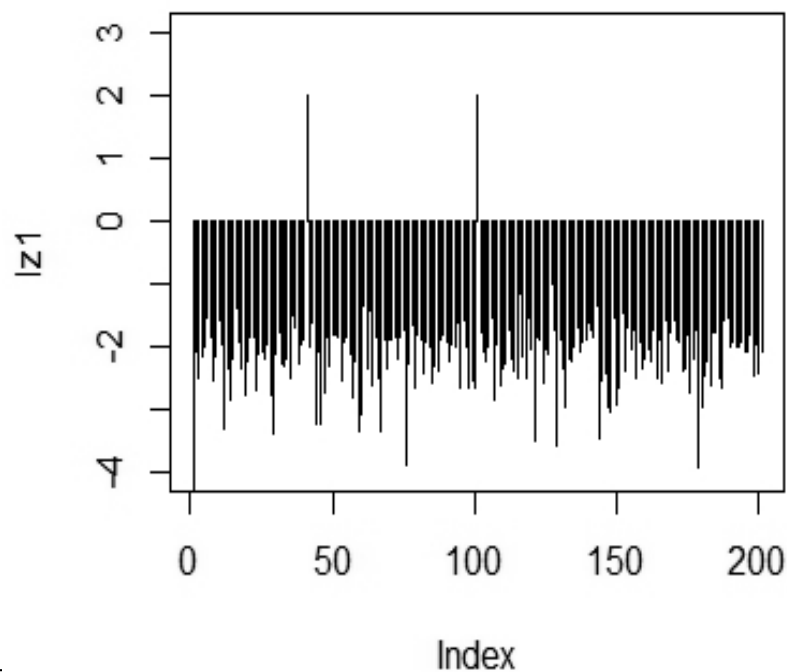
```
z <- period(y,window=0)
```

```
z1 <- z$period
```

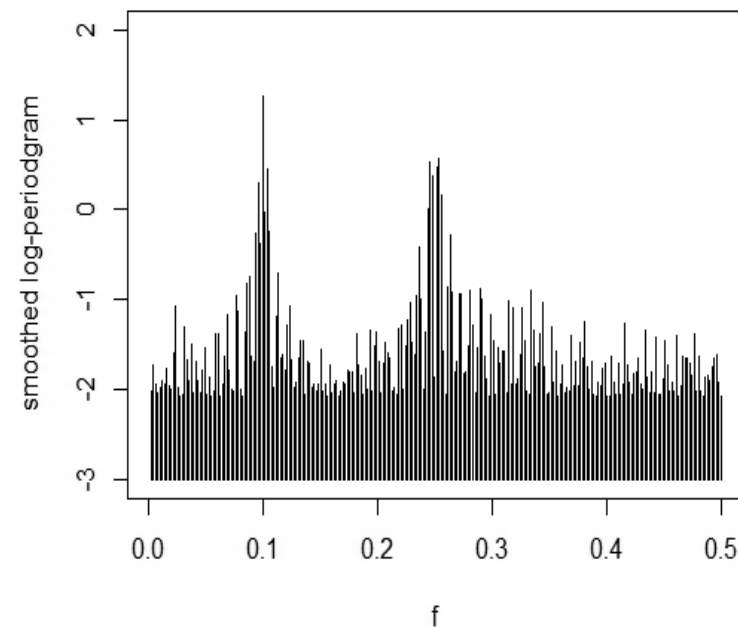
```
lz1 <- log10(z1)
```

```
plot(lz1,type="h")
```

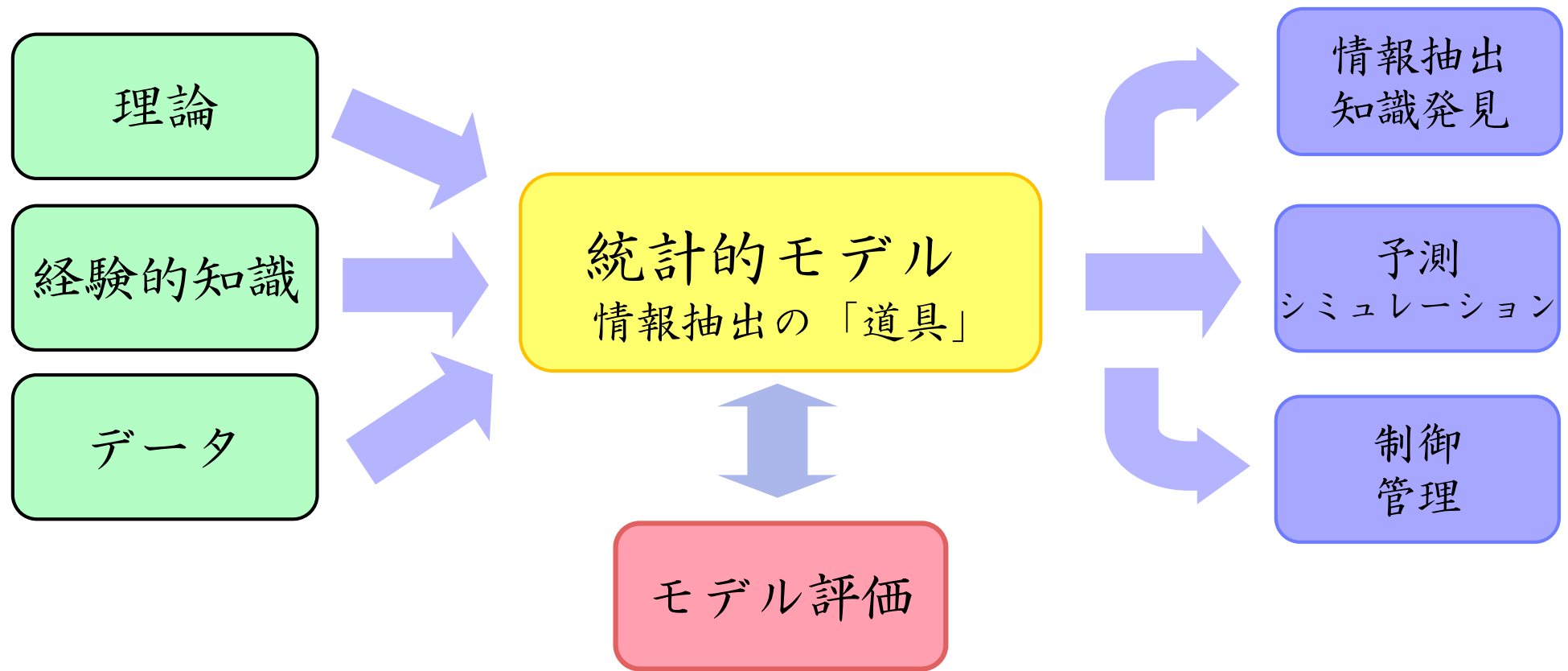
ピリオドグラム



FFTピリオドグラム

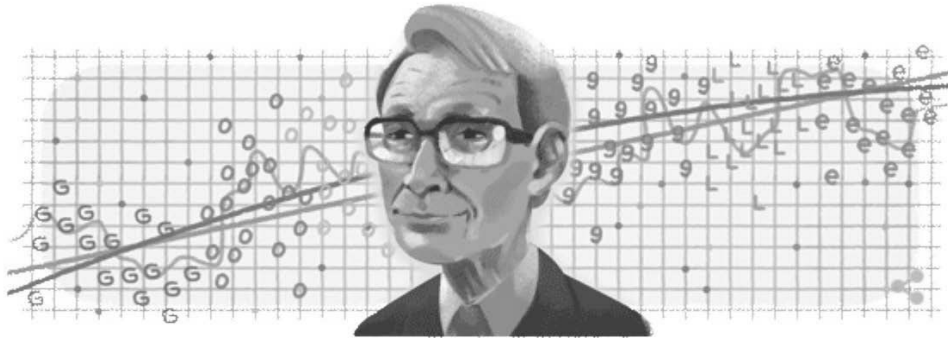


統計的モデリングと評価



- 統計的推論の結果はモデルに依存する
- モデル評価・選択が重要

H. Akaike Became Google's Top Logo 11/6/2017



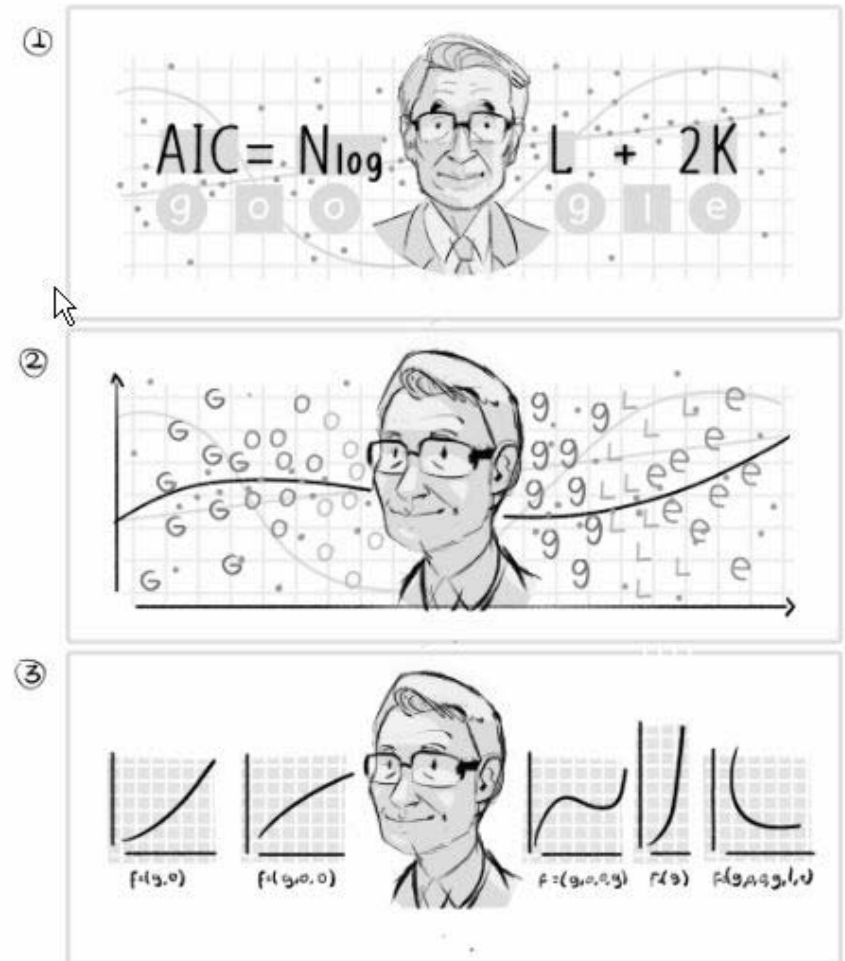
If you've ever conducted a statistical analysis, you might've spent hours thinking about which variables to include and the impact each would have on the outcome. But to ensure the model itself is accurate, shouldn't someone measure the measurers?

In the early 1950s, a young Japanese scientist named Hirotugu Akaike asked this simple but crucial question. More than two decades of research later, he presented the answer as a simple equation known as the Akaike Information Criterion. With AIC, analysts select a model from a set of options by measuring how close the results are to the (hypothetical) truth.

For Dr. Akaike, experience was core to creativity. To get 'a direct feel of random vibrations,' for example, he bought a scooter and rode it around Mount Fuji. This first-hand experience helped him differentiate between the vibrations of riding on normal and heavily-trucked roads.

Today's Doodle portrays Dr. Akaike against a Google-inspired approximation of functions, parameters, and their respective curves.

Below are a few initial conceptualizations of the Doodle.



K-L情報量とエントロピー

$$B(g; f) = -I(g; f) = \sum g_i \log \frac{g_i}{f_i}$$

Boltzmannのエントロピー

モデル $f = (f_1, \dots, f_k)$
 n 個の独立な観測値 $(n_1, \dots, n_k)$ $n_1 + \dots + n_k = n$
相対度数 $(g_1, \dots, g_k)$
 $g_i = n_i/n$ $(n_1, \dots, n_k)$ が得られる確率

$$B(g; f) \sim \frac{1}{n} \log W$$

W: 想定したモデルから得られたサンプルの相対度数が真の分布と一致する確率

$$W = \frac{n!}{n_1! \cdots n_k!} f_1^{n_1} \cdots f_k^{n_k}$$

$$\log W! = \log n! - \sum_{i=1}^k \log n_i! + \sum_{i=1}^k n_i \log f_i$$

$$\sim n \log n - n - \sum_{i=1}^k n_i \log n_i + \sum_{i=1}^k n_i + \sum_{i=1}^k n_i \log f_i$$

$$= - \sum_{i=1}^k n_i \log \frac{n_i}{n} + \sum_{i=1}^k n_i \log f_i$$

$$= \sum_{i=1}^k n_i \log \frac{f_i}{g_i}$$

$$= n \sum_{i=1}^k g_i \log \frac{f_i}{g_i} = nB(g; f)$$

数値積分によるKL情報量の計算

```
# g:gauss, f:gauss
```

```
klinfo(1, c(0, 1), 1, c(0.1, 1.5), 8)
```

x_k	k	dx	KLI	gint
8.00	16	1.0000	0.03939926	1.000000001
8.00	32	0.5000	0.03939922	1.000000000
8.00	64	0.2500	0.03939922	1.000000000
8.00	128	0.1250	0.03939922	1.000000000

```
# g:gauss, f:cauchy
```

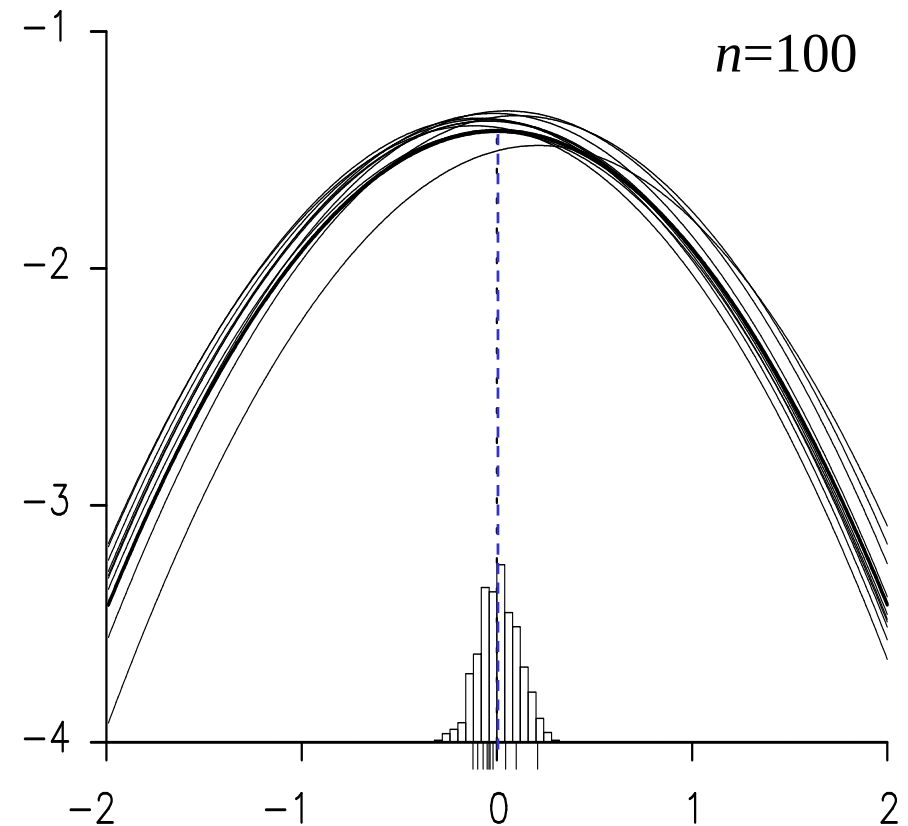
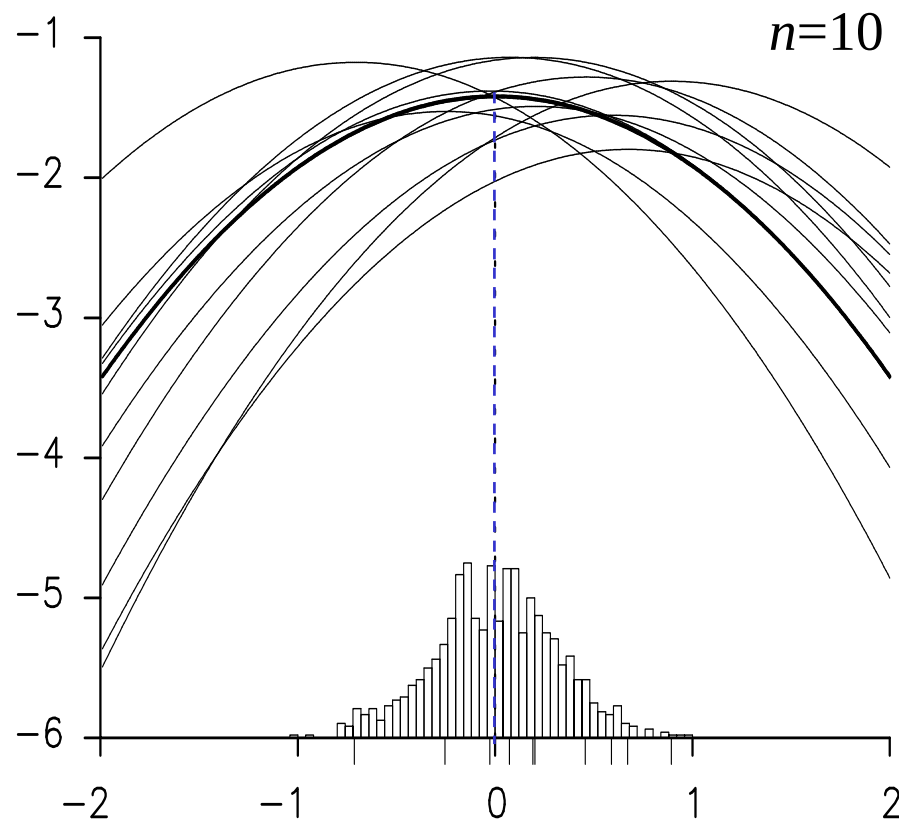
```
klinfo(1, c(0, 1), 2, c(0, 1), 8)
```

x_k	k	dx	KLI	gint
8.00	16	1.0000	0.25620181	1.000000001
8.00	32	0.5000	0.25924202	1.000000000
8.00	64	0.2500	0.25924453	1.000000000
8.00	128	0.1250	0.25924453	1.000000000

最尤推定値の例(正規分布の平均)

$$y \sim N(0,1), \quad \text{model: } N(\mu,1)$$

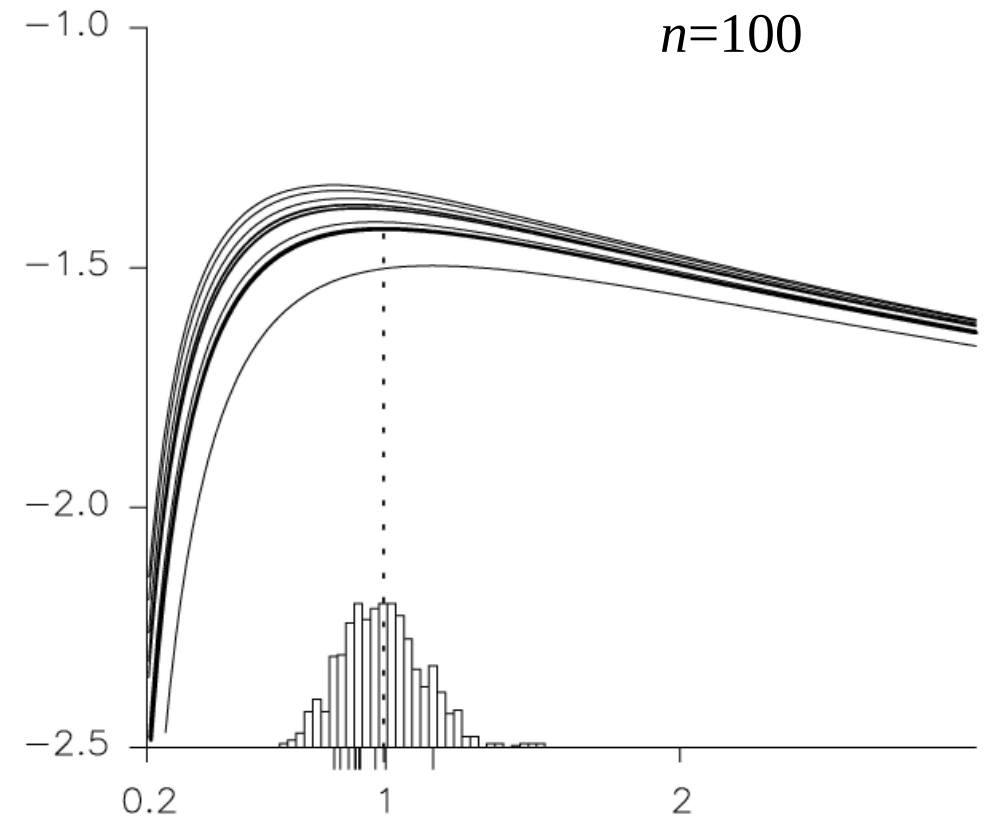
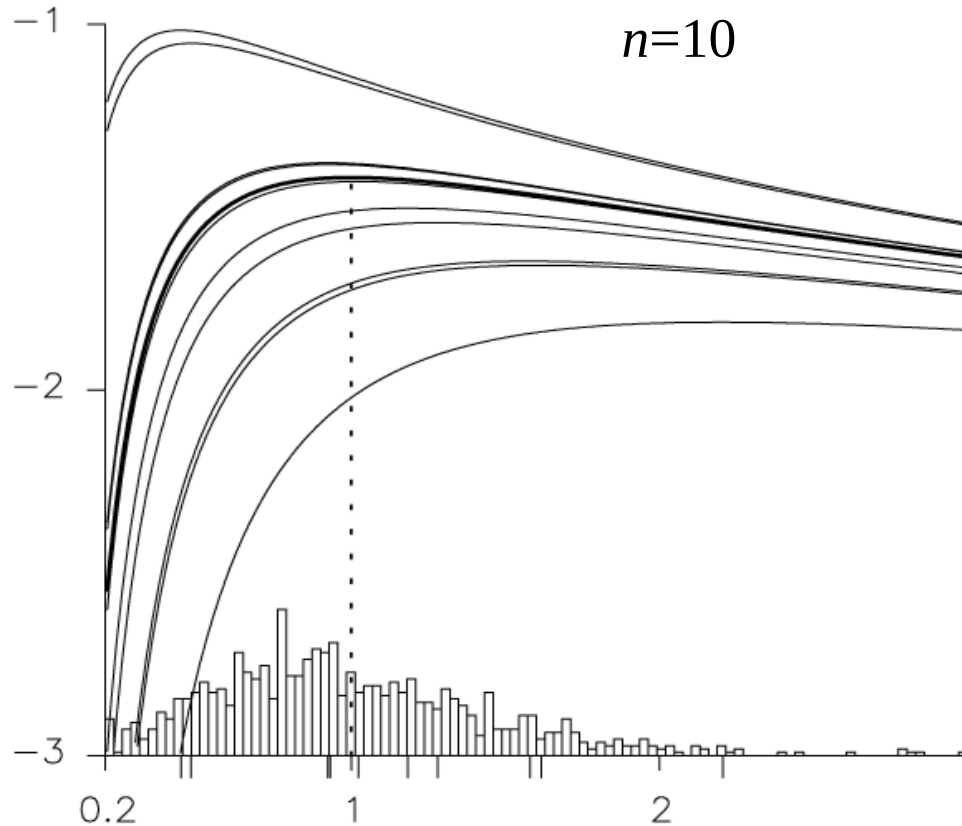
———— Expected log-likelihood
———— log-likelihood



最尤推定値の例（正規分布の分散）

$$y \sim N(0,1), \quad \text{model: } N(0, \sigma^2)$$

— Expected log-likelihood
— log-likelihood

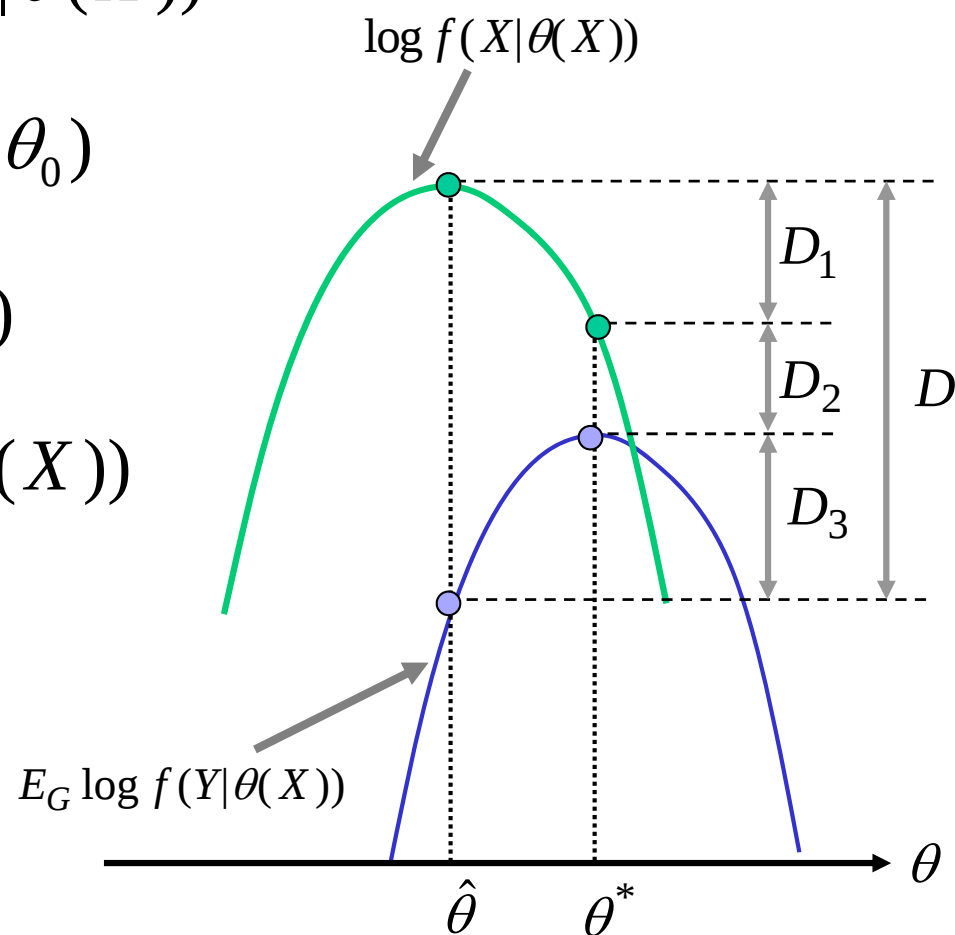


バイアスの評価

対数尤度

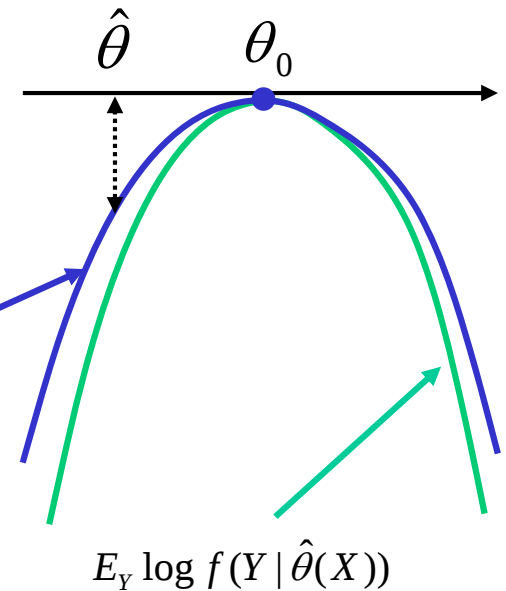
平均対数尤度

$$\begin{aligned} D &= \frac{1}{n} \log f(X | \hat{\theta}(X)) - E_Y \log f(Y | \hat{\theta}(X)) \\ &= \frac{1}{n} \log f(X | \hat{\theta}(X)) - \frac{1}{n} \log f(X | \theta_0) \\ &\quad + \frac{1}{n} \log f(X | \theta_0) - E_Y \log f(Y | \theta_0) \\ &\quad + E_Y \log f(Y | \theta_0) - E_Y \log f(Y | \hat{\theta}(X)) \\ &= D_1 + D_2 + D_3 \end{aligned}$$



D_3, D_1, D_2 の評価

$$\begin{aligned}
 E_Y \log f(Y | \hat{\theta}(X)) &= 0 \\
 &\approx E_Y \log f(Y | \theta_0) + \frac{\partial}{\partial \theta} E_Y \log f(X | \theta_0)(\hat{\theta} - \theta_0) \\
 &= + \frac{1}{2}(\hat{\theta} - \theta_0)^T \boxed{E_Y \frac{\partial^2}{\partial \theta \partial \theta} \log f(Y | \theta_0)}(\hat{\theta} - \theta_0) \\
 &= E_Y \log f(Y | \theta_0) - \frac{1}{2}(\hat{\theta} - \theta_0)^T \boxed{J(\theta_0)}(\hat{\theta} - \theta_0)
 \end{aligned}$$



$$E_X \{(\hat{\theta} - \theta_0)^T J(\theta_0)(\hat{\theta} - \theta_0)\} = \frac{1}{n} \text{tr} \{I(\theta_0) J^{-1}(\theta_0)\}$$

$$E_X \{D_3\} = E_X \{E_Y \log f(Y | \theta_0) - E_Y \log f(Y | \hat{\theta}(X))\} \approx \frac{1}{2n} \text{tr} \{IJ^{-1}\}$$

$$\begin{aligned}
 E_X \{D_1\} &= E_X \{\log f(X | \hat{\theta}(X)) - \log f(X | \theta_0)\} \approx \frac{1}{2n} \text{tr} \{IJ^{-1}\} \\
 E_X \{D_2\} &= E_X \left\{ \frac{1}{n} \log f(X | \theta_0) - E_Y \log f(Y | \theta_0) \right\} = 0
 \end{aligned}$$

$$b(G) = E_X \{ D \} = E_X \{ D_1 + D_2 + D_3 \}$$

$$b_{\text{TIC}}(G) = \text{tr} \{ I(G) J(G)^{-1} \}$$

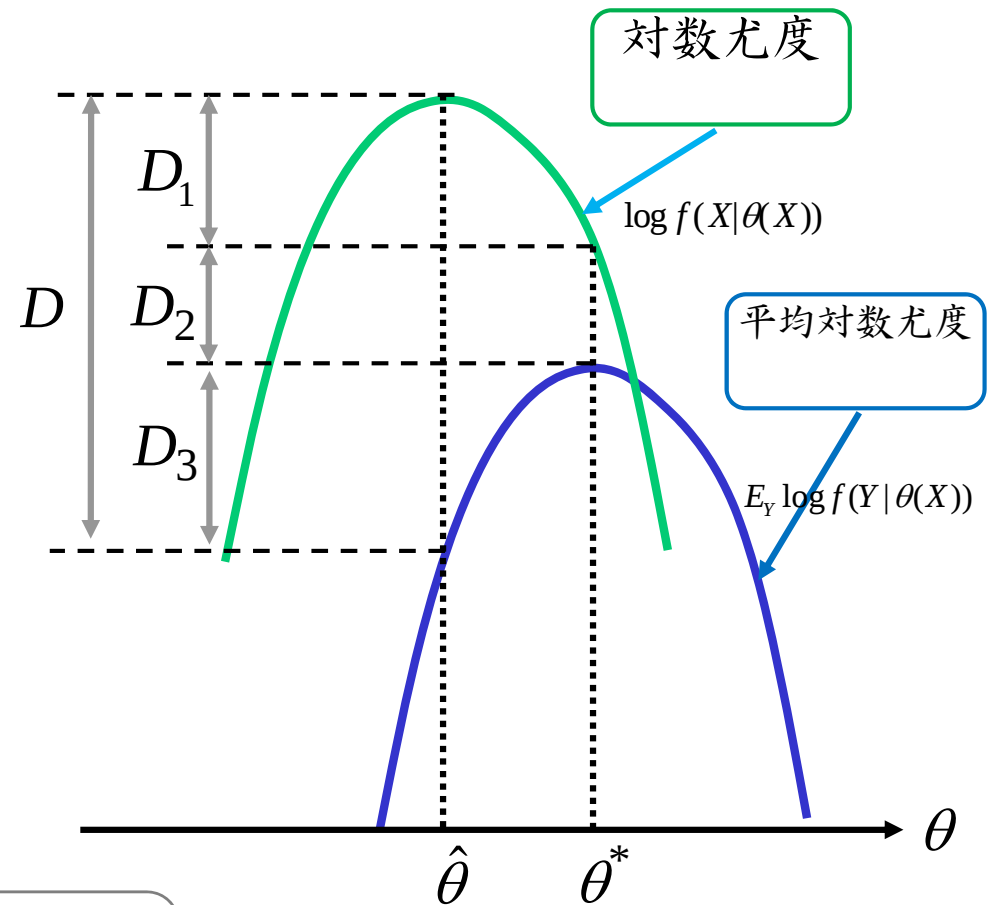
$$I(G) = E_X \left[\frac{\partial \log f(X | \theta)}{\partial \theta} \frac{\partial \log f(X | \theta)}{\partial \theta'} \right]$$

Fisher情報量

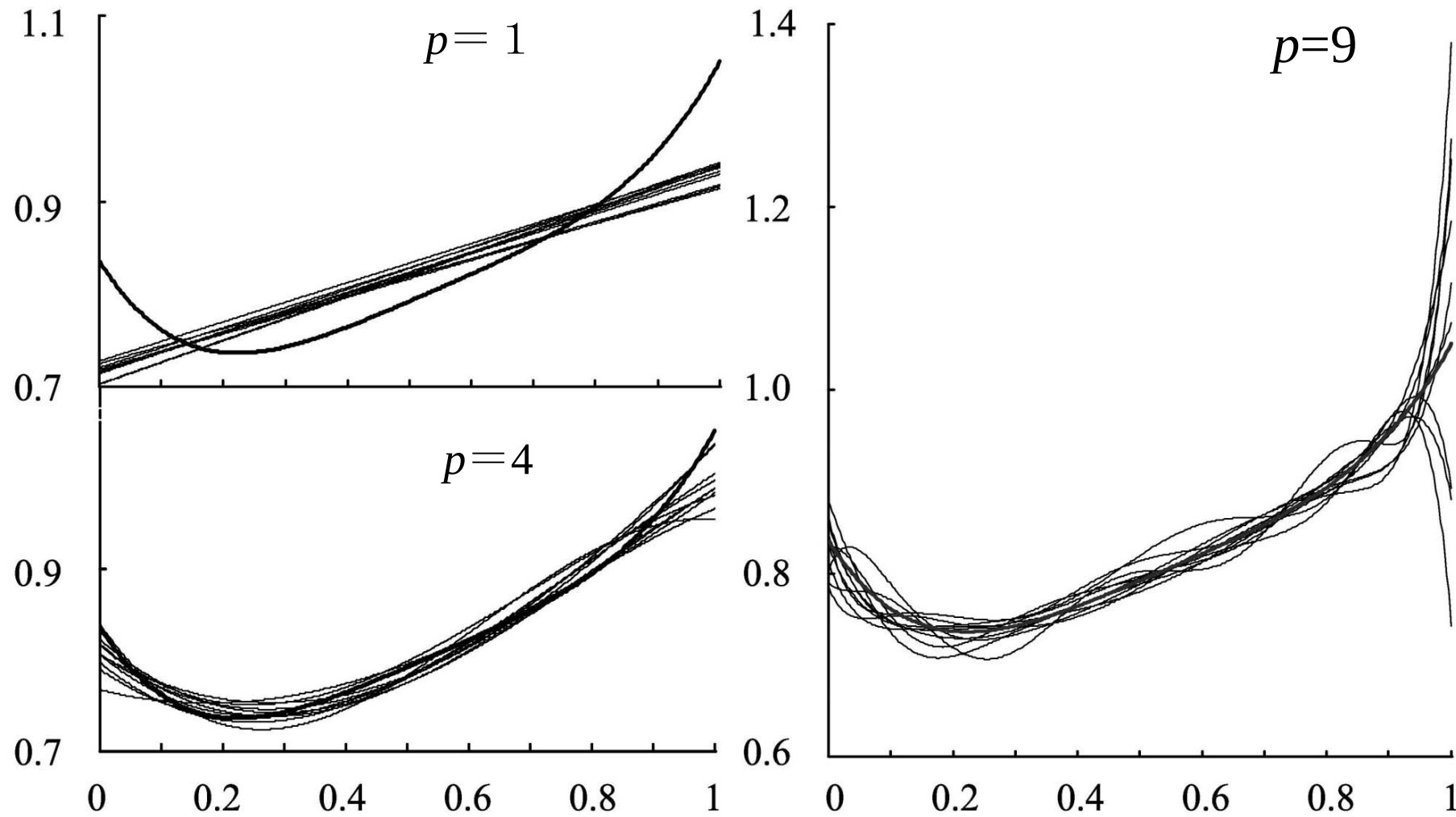
$$J(G) = -E_X \left[\frac{\partial^2 \log f(X | \theta)}{\partial^2 \theta} \right]$$

ヘッセ行列の期待値

$$\text{IC} = -2 \log f(x | \hat{\theta}) + 2 \hat{b}(G)$$



モデル選択例：多項式回帰の次数



Selection of the Bin Size of a Histogram

$$P(\{n_j\}|\{p_j\}) = \frac{n!}{n_1! \cdots n_k!} p_1^{n_1} \cdots p_k^{n_k}$$

$$\ell(p_1, \dots, p_k) = C + \sum_{j=1}^k n_j \log p_j$$

$$\hat{p}_j = \frac{n_j}{n}$$

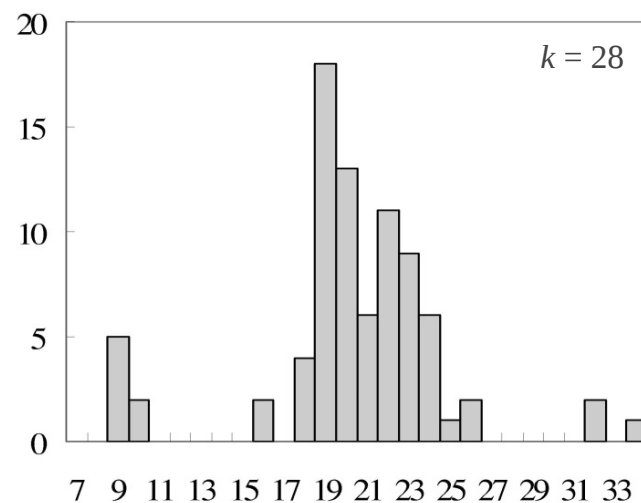
$$\text{AIC}_k = (-2) \left\{ C + \sum_{j=1}^k n_j \log \left(\frac{n_j}{j} \right) \right\} + 2(k-1)$$

Galaxy data (Roeder (1990))

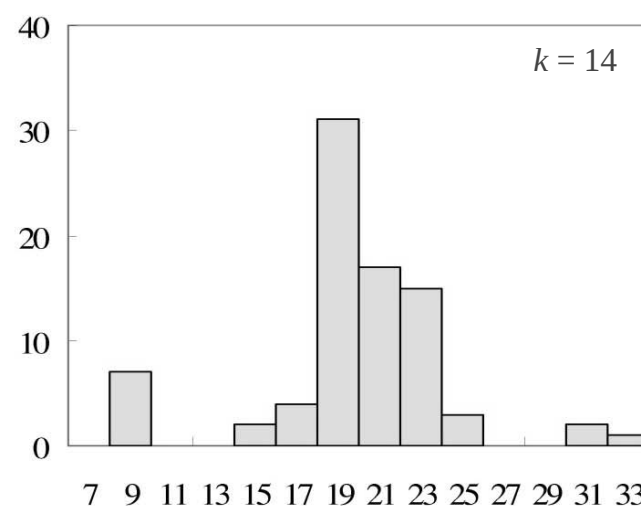
0 5 2 0 0 0 0 0 2 0 4 18 13 6
11 9 6 1 2 0 0 0 0 0 2 0 1 0

Bin Size	log-LK	AIC
28	-189.19	432.38
14	-197.72	421.43
7	-209.52	431.03

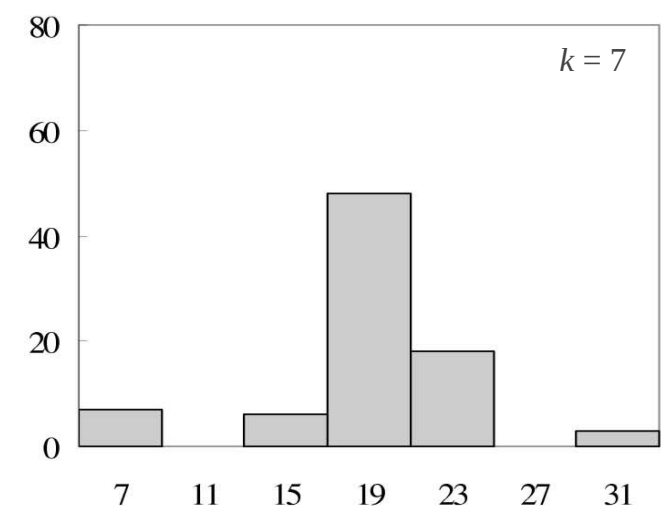
Histogram of galaxy data



Best



Too small



Variable Selection for a Regression Model

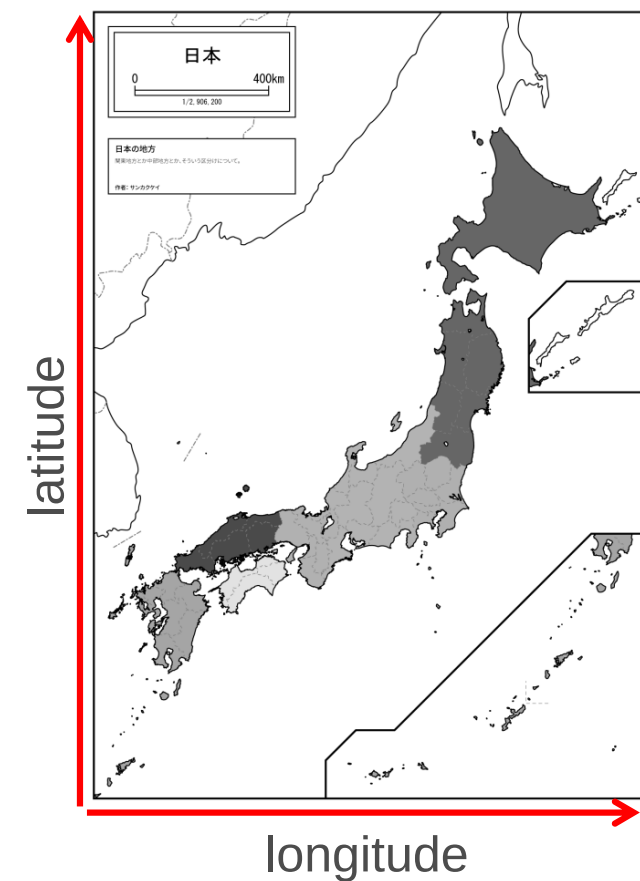
y_n : Temperature, x_{1n} : Latitude, x_{2n} : Longitude, x_{3n} : Altitude

$$y_n = a_0 + a_1 x_{1n} + a_2 x_{2n} + a_3 x_{3n} + \varepsilon_n, \quad \varepsilon_n \sim N(0, \sigma^2)$$

Select variables among x_1, x_2, x_3 appropriate to predict y_n

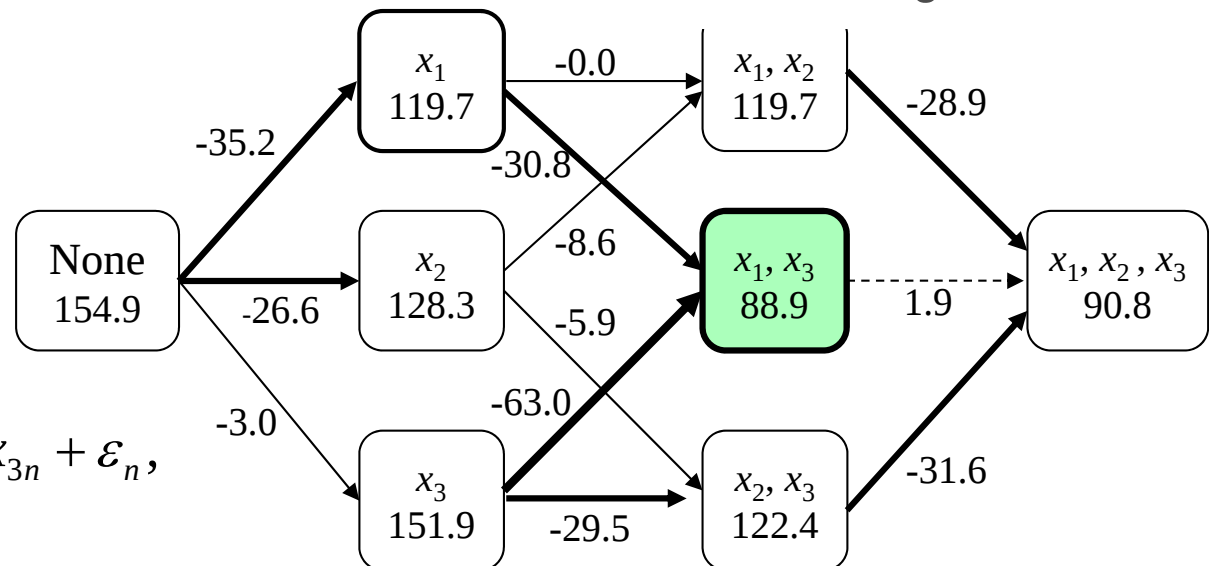
$$\ell(a_0, a_1, a_2, a_3, \sigma^2) = -\frac{n}{2} \log(2\pi\sigma^2) - \frac{1}{2\sigma^2} \sum_{i=1}^n (y_n - a_0 - \sum_{j=1}^k a_j x_{jn})^2$$

$$\text{AIC}(x_{1n}, x_{2n}, x_{3n}) = n(\log 2\pi + 1) + n \log \hat{\sigma}^2 + 2(k + 2)$$



Selected model

$$y_n = 40.490 - 1.208x_{1n} - 0.010x_{3n} + \varepsilon_n, \\ \varepsilon_n \sim N(0, 1.490)$$



モデル選択例：分布の形状の選択

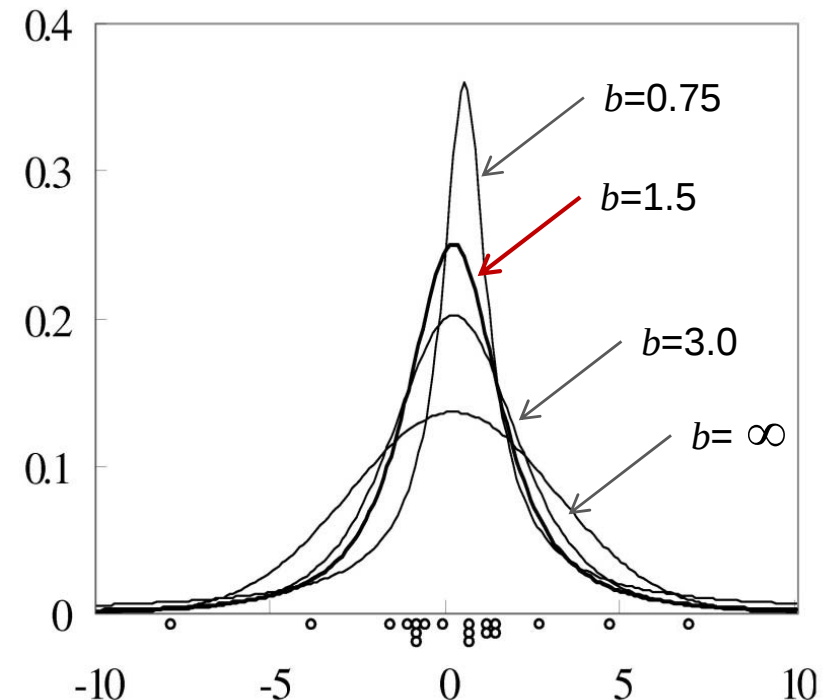
$$f(y | \mu, \tau^2, b) = \frac{C}{(y^2 + \tau^2)^b}$$

Pearson's family of distributions
Select the shape parameter b

$$\ell(\mu, \tau^2, b) = \sum_{n=1}^N \log f(y_n | \mu, \tau^2, b)$$

$$= N \left\{ (b - \frac{1}{2}) \log \tau^2 + \log \Gamma(b) - \log(b - \frac{1}{2}) - \log \Gamma(\frac{1}{2}) \right\} - b \sum_{n=1}^N \log \{(y_n - \mu)^2 + \tau^2\}$$

b	μ	τ^2	Log-L	AIC
0.60	0.801	0.030	-58.84	121.69
0.75	0.506	0.431	-51.40	106.79
1.00	0.189	1.380	-47.87	99.73
1.50	0.185	4.152	-47.07	98.14
2.00	0.201	8.395	-47.43	98.86
2.50	0.214	13.87	-47.82	99.63
3.00	0.222	20.21	-48.12	100.25
∞	0.166	8.545	-49.83	103.66



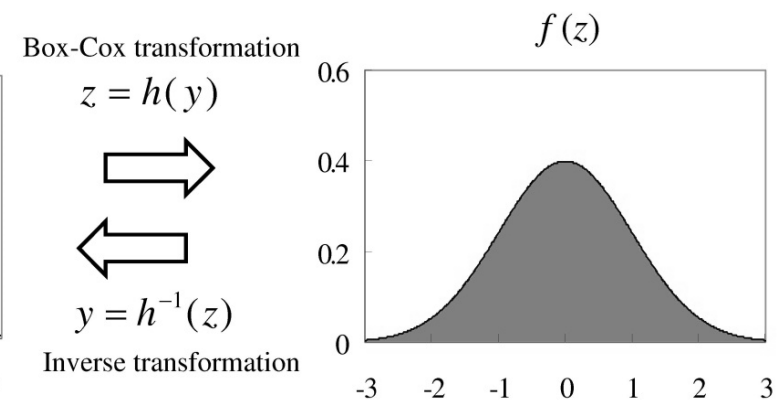
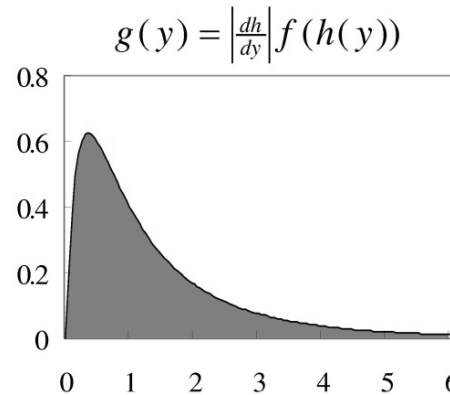
Selection of the Box-Cox transformation

$$z_n = h(y_n) = \begin{cases} \lambda^{-1}(y_n^\lambda - 1) & \text{for } \lambda \neq 0 \\ \log y_n & \text{for } \lambda = 0 \end{cases}$$

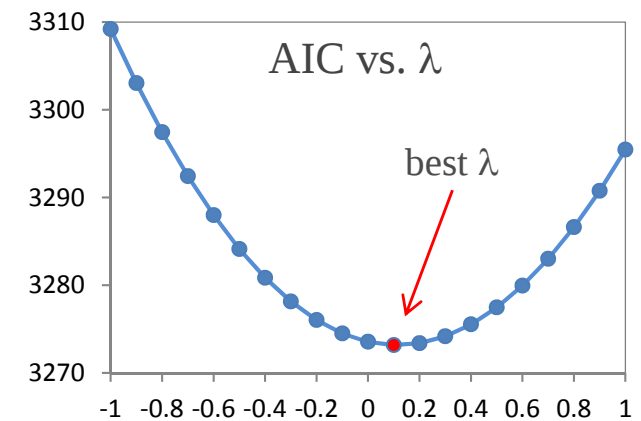
$$g(y) = \left| \frac{dh_\lambda}{dy} \right| f(h(y))$$

$$AIC'_z = AIC_z - 2 \log \left| \frac{dh_\lambda}{dy} \right|$$

Jacobian of the transformation



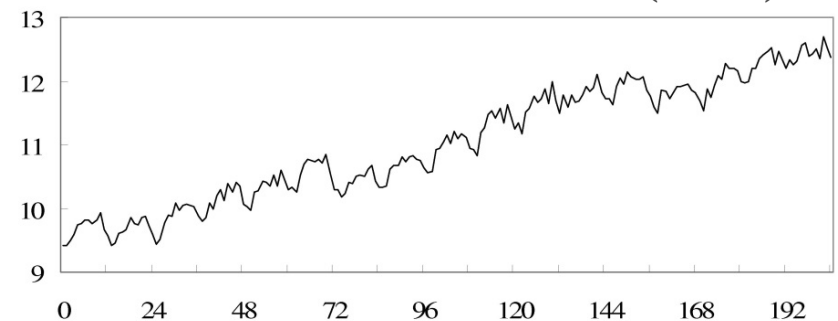
λ	-1.0	-0.5	0.0	0.1	0.2	0.5	1.0
log-L	1374.9	630.5	-121.1	-272.2	-423.7	-879.9	-1645.3
AIC	-2745.7	-1257.0	246.1	548.5	854.4	1763.8	3295.5
AIC'	3309.2	3284.1	3273.6	3273.2	3273.4	3277.5	3295.5



Original WHARD data (US BLS)



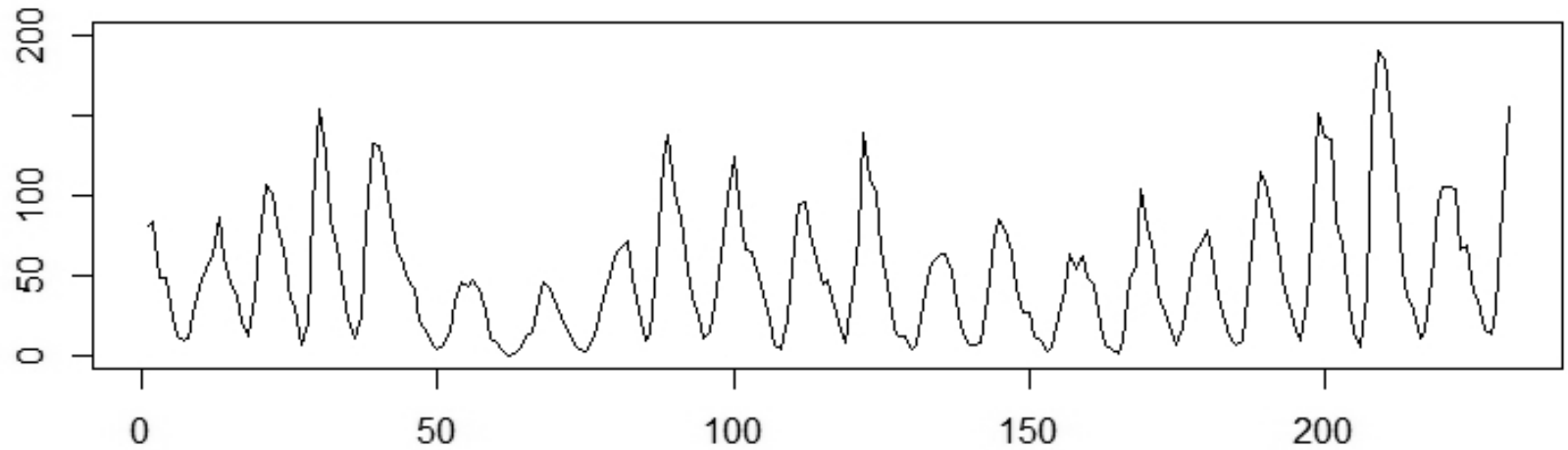
Best Box-Cox transformation ($\lambda=0.1$)



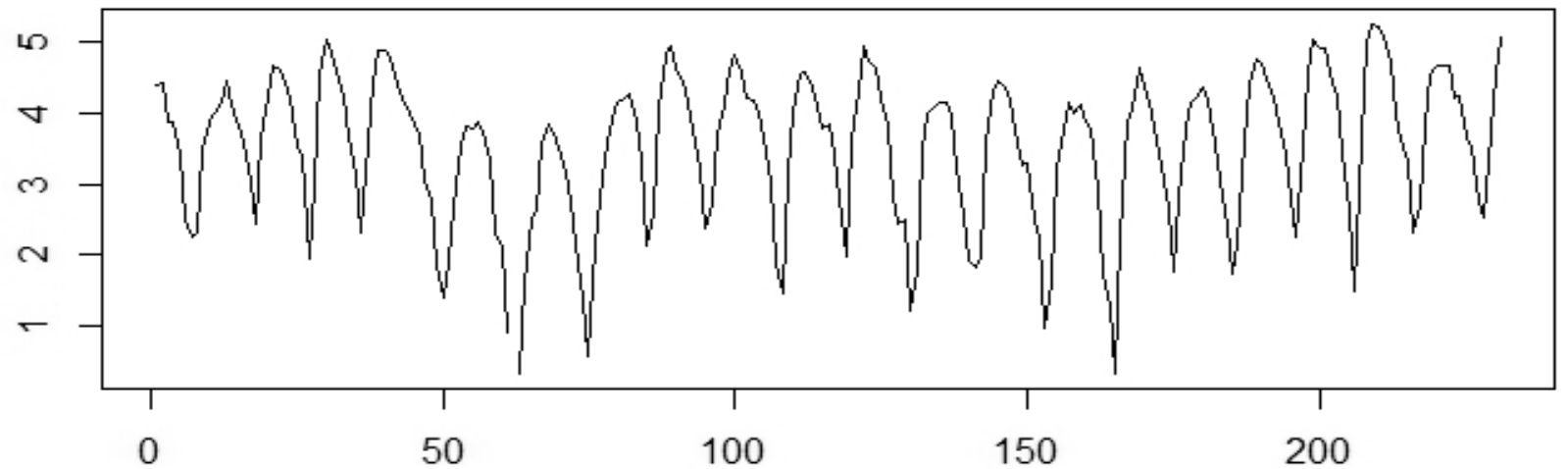
太陽黒点数データ

```
plot(sunspot,ylim=c(0,200))  
y <- log( sunspot )  
plot(y)
```

$y(n)$



$\log y(n)$



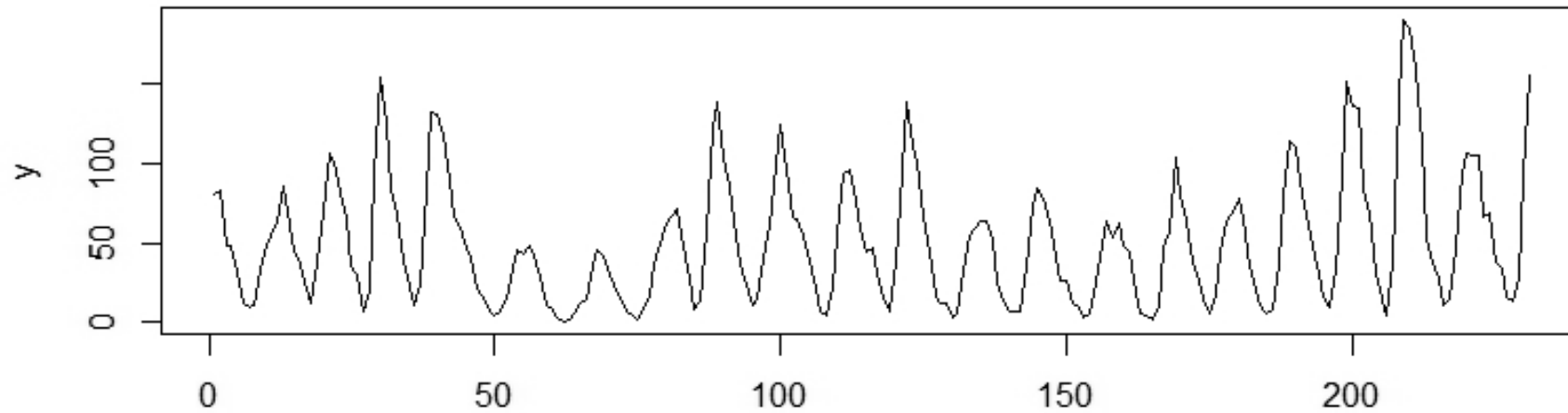
data(Sunspot) # Sun spot number data
 boxcox(Sunspot)

lambda	aic'	LL'	aic	LL	mean	variance
1.00	2360.26	-1178.13	2360.26	-1178.13	4.909502e+01	1.575552e+03
0.90	2335.22	-1165.61	2174.47	-1085.24	3.545844e+01	7.049401e+02
0.80	2313.48	-1154.74	1991.98	-993.99	2.591126e+01	3.199262e+02
0.70	2295.33	-1145.66	1813.07	-904.54	1.917397e+01	1.474669e+02
0.60	2281.11	-1138.56	1638.11	-817.05	1.437922e+01	6.914276e+01
0.50	2271.26	-1133.63	1467.50	-731.75	1.093610e+01	3.303737e+01
0.40	2266.32	-1131.16	1301.81	-648.91	8.439901e+00	1.612487e+01
0.30	2267.05	-1131.52	1141.79	-568.90	6.611858e+00	8.065706e+00
0.20	2274.59	-1135.29	988.58	-492.29	5.258840e+00	4.155209e+00
0.10	2290.79	-1143.40	844.03	-420.01	4.246205e+00	2.222464e+00
0.00	2318.78	-1157.39	711.27	-353.63	3.479466e+00	1.250918e+00
-0.10	2363.66	-1179.83	595.39	-295.70	2.891856e+00	7.574966e-01
-0.20	2432.86	-1214.43	503.84	-249.92	2.435839e+00	5.096385e-01
-0.30	2534.61	-1265.31	444.85	-220.42	2.077302e+00	3.947690e-01
-0.40	2673.75	-1334.88	423.23	-209.62	1.791544e+00	3.595107e-01
-0.50	2848.16	-1422.08	436.89	-216.45	1.560501e+00	3.814048e-01
-0.60	3050.32	-1523.16	478.30	-237.15	1.370809e+00	4.562814e-01
-0.70	3271.90	-1633.95	539.12	-267.56	1.212437e+00	5.937308e-01
-0.80	3506.54	-1751.27	613.01	-304.51	1.077716e+00	8.175441e-01
-0.90	3750.16	-1873.08	695.88	-345.94	9.606427e-01	1.170321e+00
-1.00	4000.25	-1998.13	785.23	-390.61	8.563591e-01	1.722986e+00

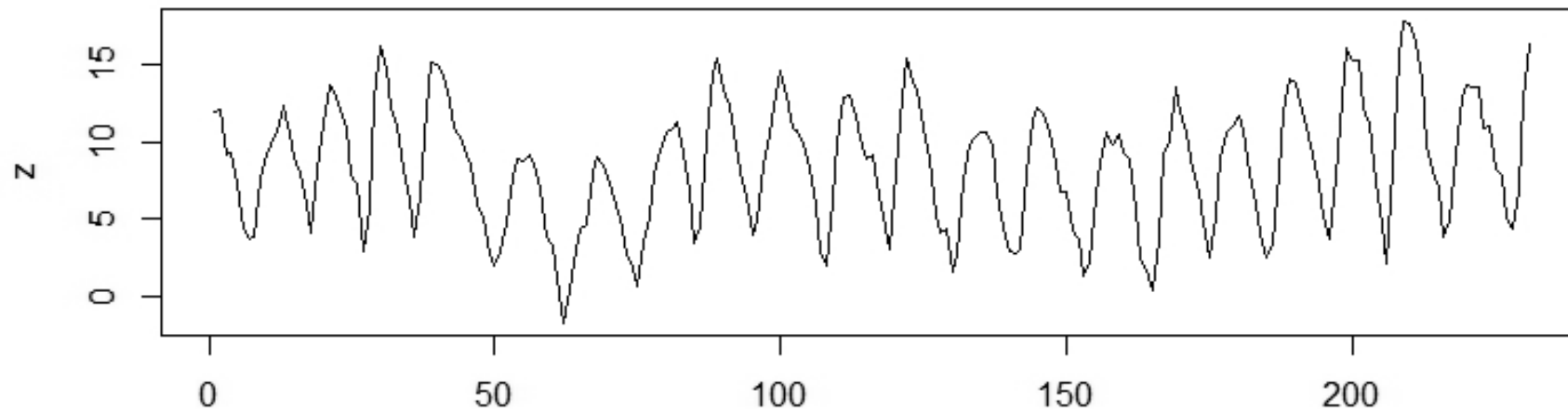
lambda = 0.40 AIC' minimum = 2266.32

```
data(Sunspot) # Sun spot number data  
boxcox(Sunspot)
```

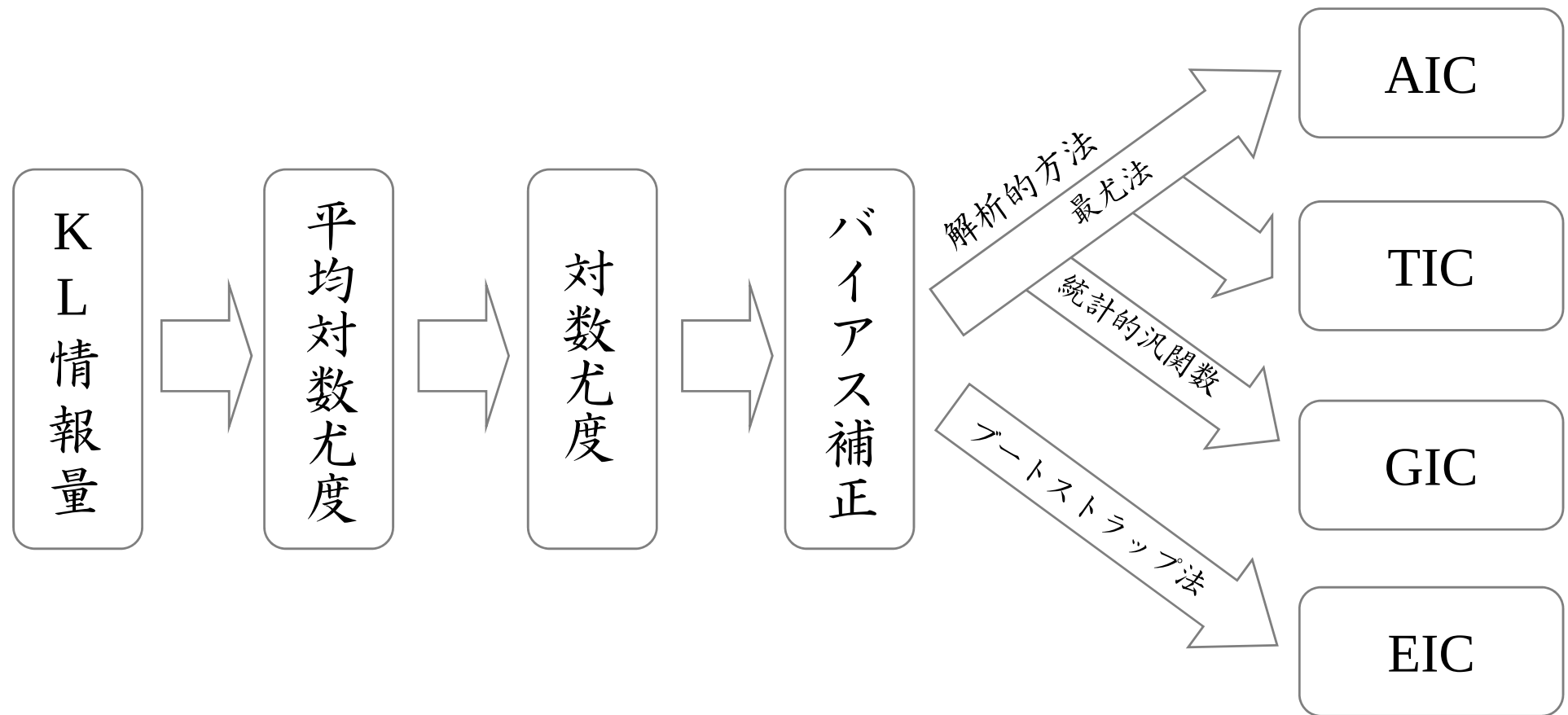
original data y



transeformed data z



情報量規準の系譜



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